

(Estimation of the magnetic field) and accretion flow geometry of Low Mass X-ray Binaries

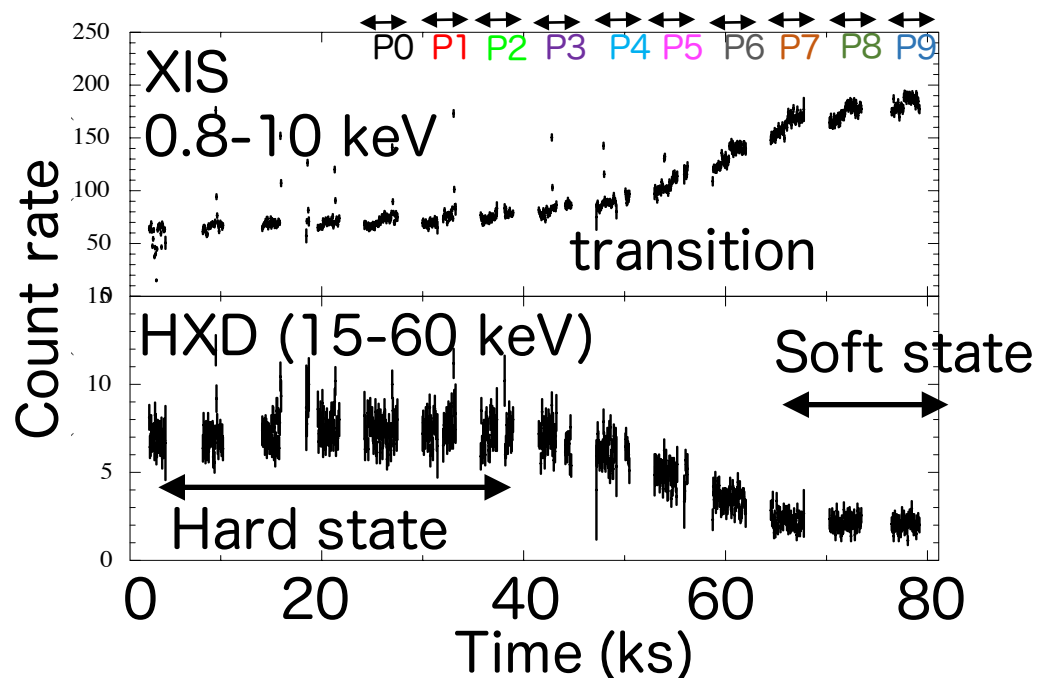
Ko Ono^[A], Kazuo Makishima^[B], Zhongli Zhang^[C],
Kazuhiro Nakazawa^[A]

[A] The University of Tokyo [B] RIKEN [C] Shanghai
Observatory

Aquila X-1 during a hard-to-soft state transition

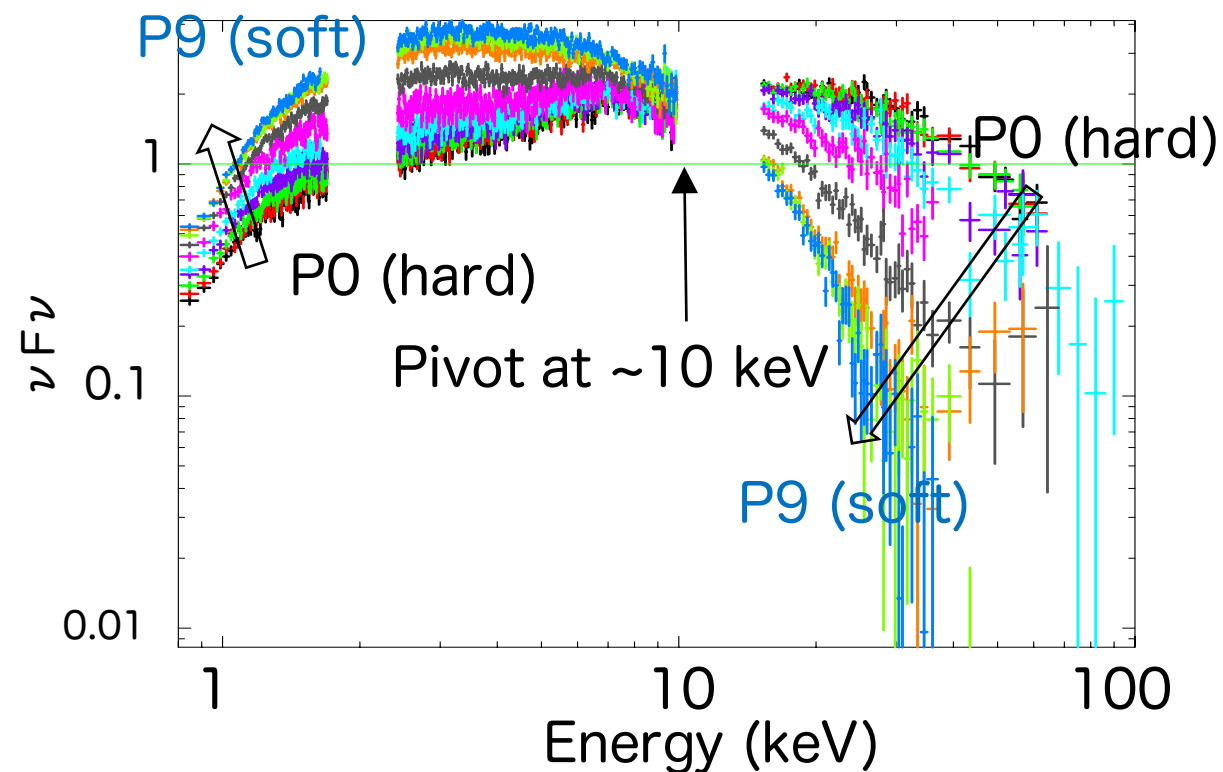
NS LMXBs:

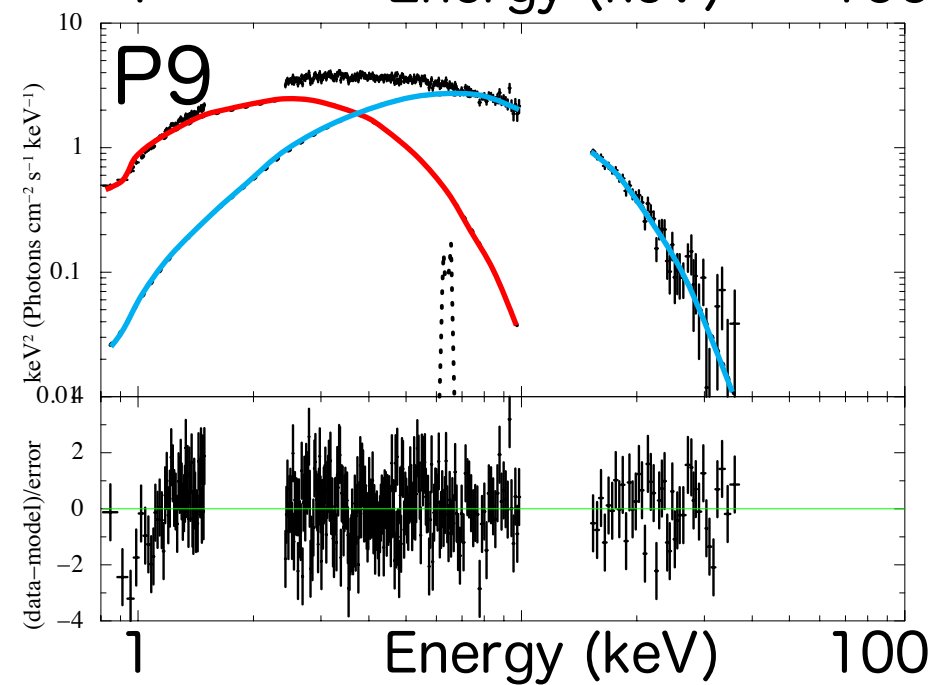
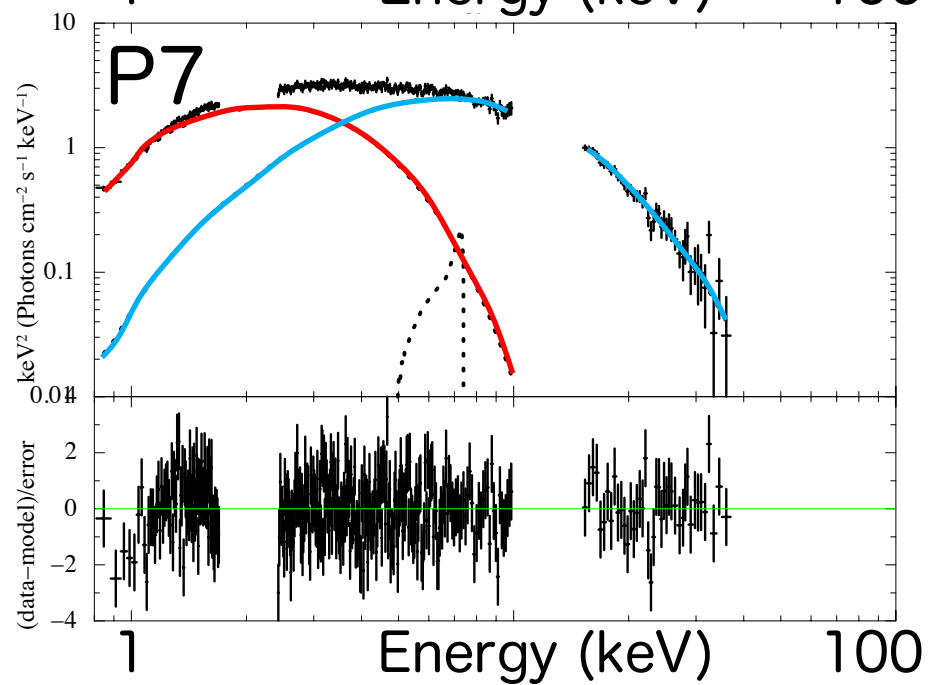
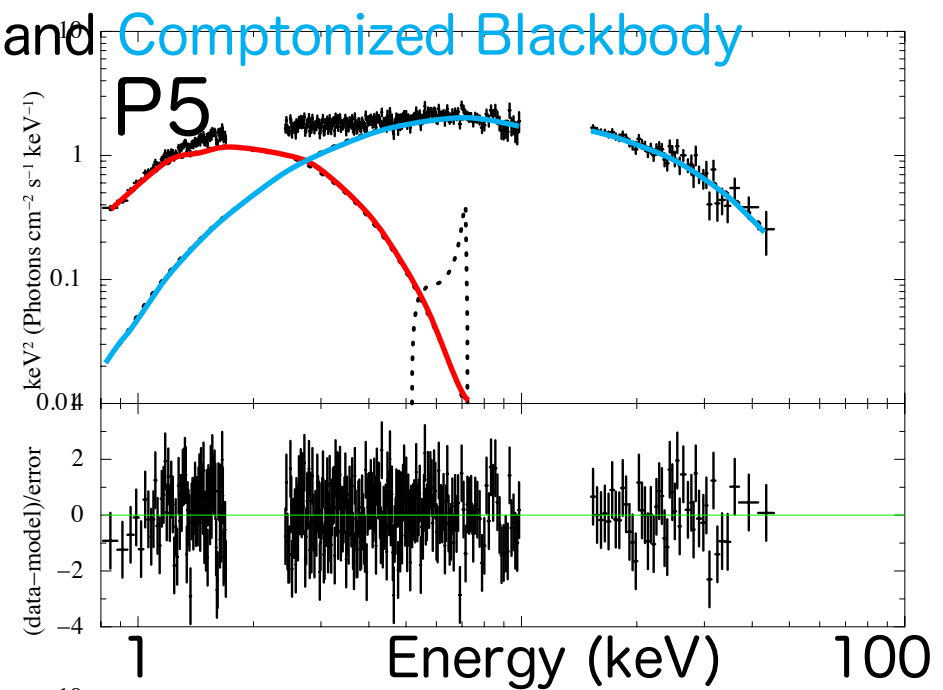
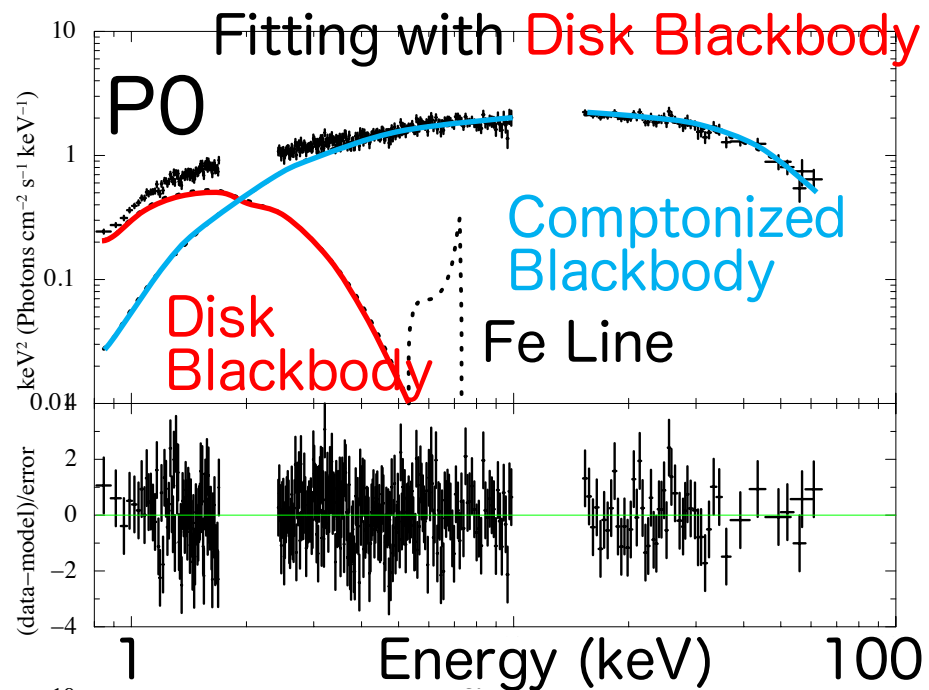
- The soft and hard state are explained by an **accretion disk**, and a **blackbody Comptonized** weakly and strongly, respectively.
- However, how about in between them?



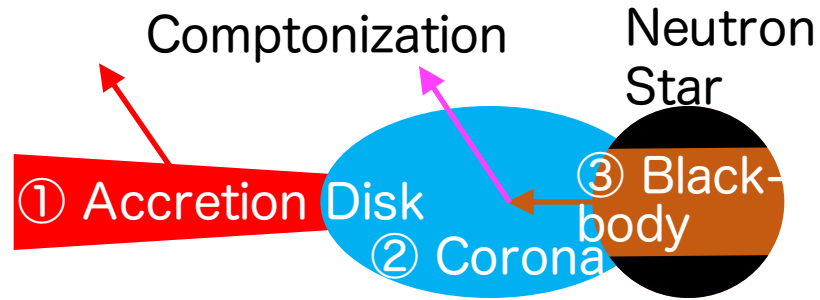
Aquila X-1:

- Recurrent transient
- Observed with Suzaku on 2011 October 21.
- Hard-to-soft state transition took place during the observation, on ~20 ksec.

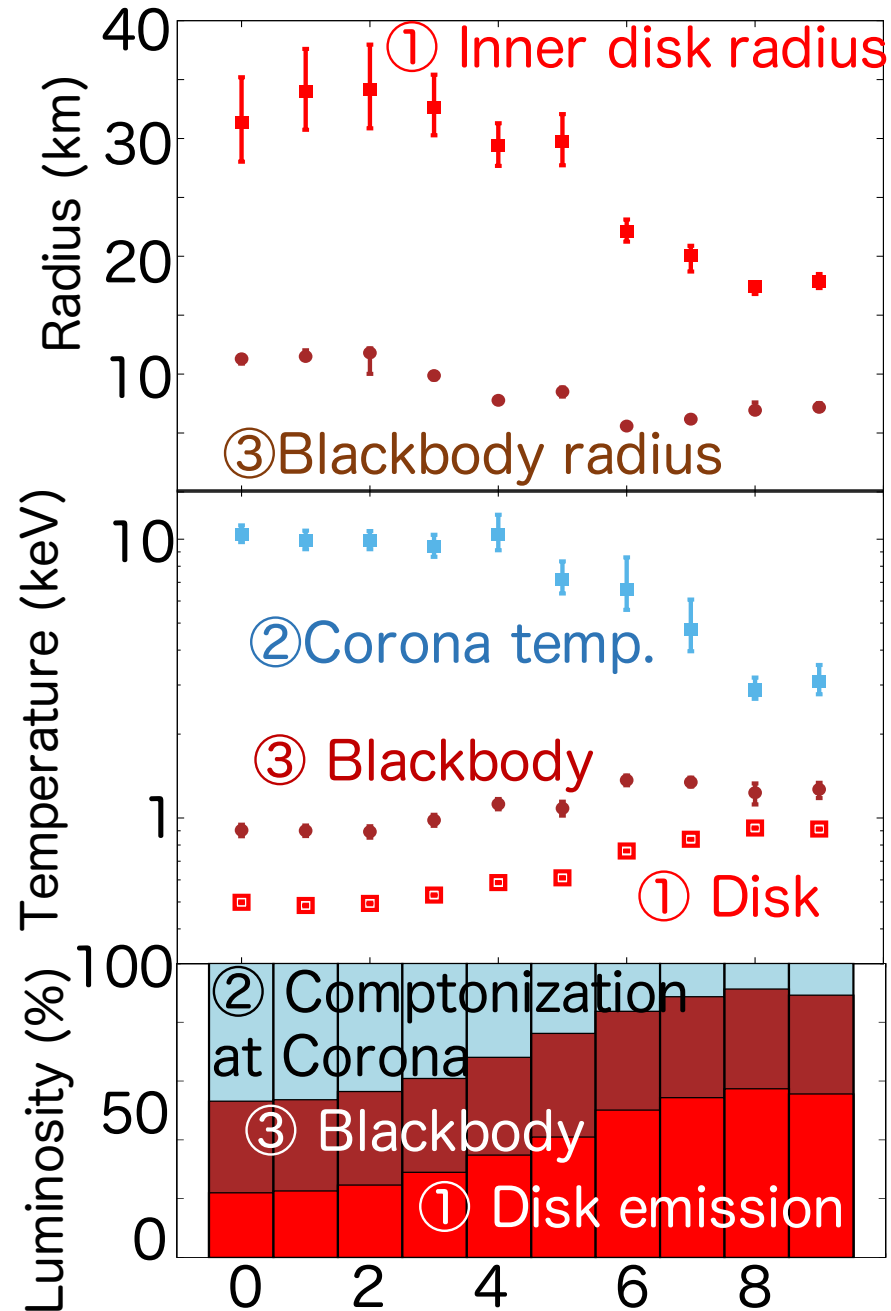




Discussion & Conclusion



- ① The opt. thick disk continues closer to the NS
- ② Corona shrinks, becomes cooler (~ 3 keV), and less luminous
- ③ Blackbody gets confined to the equator



Parameters in the two states are connected continuously.

This justifies the spectral decomposition, **Disk**, **Corona** and **Blackbody**, throughout both states.

$B < 10^8$ G, assuming $R_A < 10$ km at P0-P9