



Estimate of the Fermi LAT sensitivity for gamma-ray polarization

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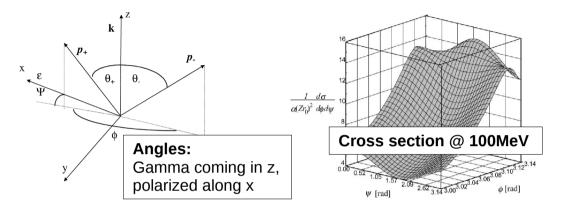


Alliance for Astroparticle Physics



Polarization with the LAT

Polarization introduce azimuthal modulation in event distribution

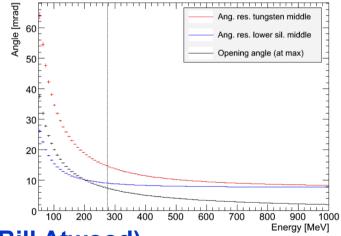


Multiple Scattering vs. Pair opening Angle:

$$\theta_{MS} = \frac{13.6 \, MeV}{E_{\gamma}/2} \sqrt{X} \left(1 + 0.038 \ln(X)\right)$$
$$\theta_{op} \approx \frac{4 \, m_e}{E_{\gamma}} \sim 1 \, deg @ 100 \, MeV$$

Azimuthal distribution: $\frac{dN}{dw} \propto (1 - A_{100} \Pi_0 \cos^2 \psi)$

$$\square$$
 · degree of linear polarization



- θ_{MS}
< θ_{op} \rightarrow silicon conversion and E<200 MeV (Bill Atwood)
- <10% of the events are usable</p>
- Measure feasible within statistic only for the brightest sources: 5sigma limit $\leftarrow \rightarrow \Pi_0 > 35\%$ (Bill Atwood, Rolf Bühler)

Gamma-ray Space Telescope



Measure systematics

Galactic latitude [deg]

180

90

0

Galactic longitude [deg]

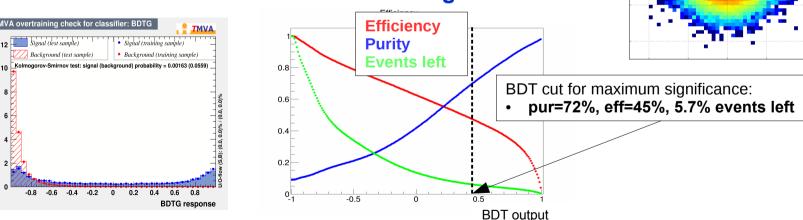
North

-90

-180

Counts (A.U

- Choose AGN stack as unpolarized source
 - Bright (~Vela)
 - Unpolarized (assuming Π_0 =60% for single AGN)
- Preliminary analysis and event selection:
 - 1-track analysis:
 - e- direction ~ best track direction
 - Need to know where $\boldsymbol{\gamma}$ comes from
 - Azimuthal angle from sky maps
 - Selection of silicon-converted events using BDT:



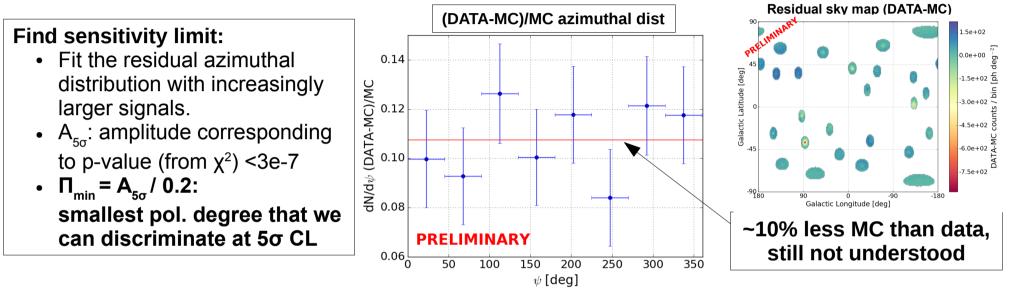
- Simulation covers 5 years:
 - Stack AGN and neighbours (>600 sources)
 - Galactic and isotropic diffuse emission

••• Stack sources ••• Auxiliary sources

Systematics estimate

Residual azimuthal distribution:

- Compatible with constant (no res. modulation) within statistic



- 5sigma sensitivity limit Π_{min} ~ 40 %
 - depends on analysis and event selection cuts.
 - Similar to purely statistical estimate: measure is limited by statistic
- Caveat:

Space Telescope

 The analysis is preliminary, need large MC simulation of polarized LAT event to develop event selection and 2-track reconstruction of azimuthal angle