

# Gas absorption and dust extinction towards the Orion Nebula Cluster

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# Motivation

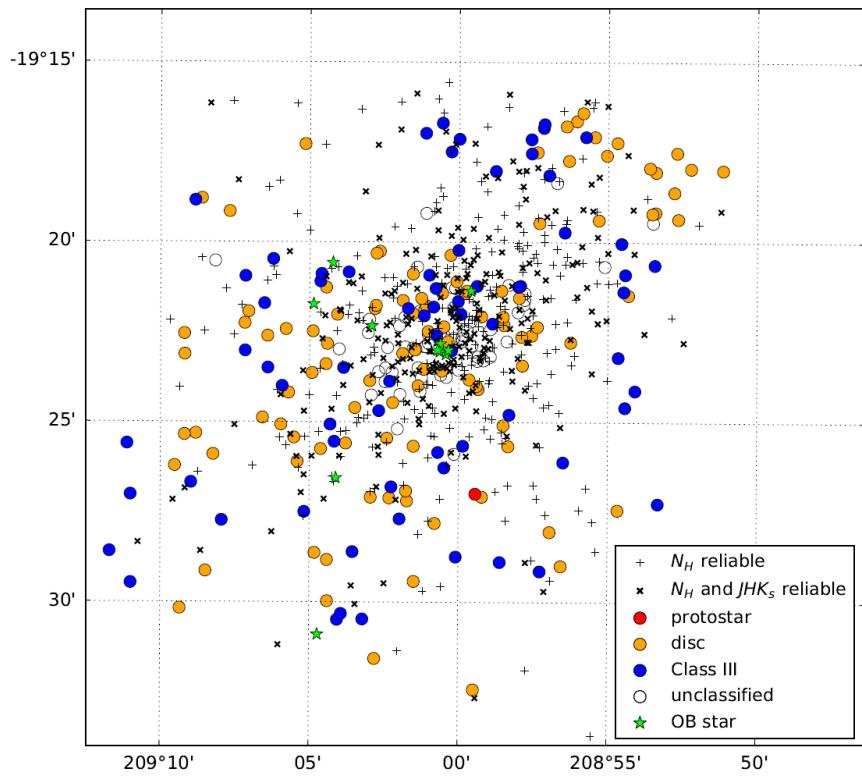
Target region	$N_H/A_V$ [cm $^{-2}$ mag $^{-1}$ ]	Reference
diffuse ISM	$2.2 \cdot 10^{21}$	Watson (2011)
	$1.8 \cdot 10^{21}$	Predehl & Schmitt (1995)
	$2.2 \cdot 10^{21}$	Gorenstein (1975)
L1641 IRAS 20050+2720 NGC 1333 Serpens $\rho$ Oph RCW 38 RCW 108 ONC	$2.2 \cdot 10^{21}$	Ryter et al. (1975) <sup>b</sup>
	$0.8 \cdot 10^{21}$	Pillitteri et al. (2013) <sup>c</sup>
	$1.2 \cdot 10^{21}$	Günther et al. (2012)
	$1.0 \cdot 10^{21}$	Winston et al. (2010)
	$0.7 \cdot 10^{21}$	Winston et al. (2007)
	$1.5 \cdot 10^{21}$	Vuong et al. (2003)
	$1.7 \cdot 10^{21}$	Winston et al. (2011)
	$2.0 \cdot 10^{21}$	Wolk et al. (2008)
	$1.4 \cdot 10^{21}$	this work

$N_H$  ... hydrogen column density

$A_V$  ... dust extinction

- A uniform  $N_H/A_V$  ratio ( $2 \cdot 10^{21} \text{ cm}^{-2} \text{ mag}^{-1}$ ) is found in the diffuse ISM throughout the sky.
- **Some star-forming regions have  $N_H/A_V$  ratios lower than the diffuse ISM, the literature gives differing explanations.**
- The ONC is a complex region, but is extensively studied: Numerous high-quality data sets provide an excellent basis for studies of the ISM.

# Methods and data



## X-ray absorption

- Fits of X-ray spectra with model of emission and absorption  
→ equivalent hydrogen column density  $N_H$
- Chandra Orion Ultradeep Project (COUP, [Getman et al. 2005](#))

## NIR extinction

- Comparison of stars' observed and intrinsic NIR colour magnitudes (NICER algorithm, [Lombardi & Alves 2001](#))  
→ V-band extinction  $A_V$
- VISTA Orion A survey (VISION, [Meingast et al. 2016](#))

## MIR magnitudes

- Classification of young stellar objects
- Spitzer Orion survey ([Megeath et al. 2012](#))

