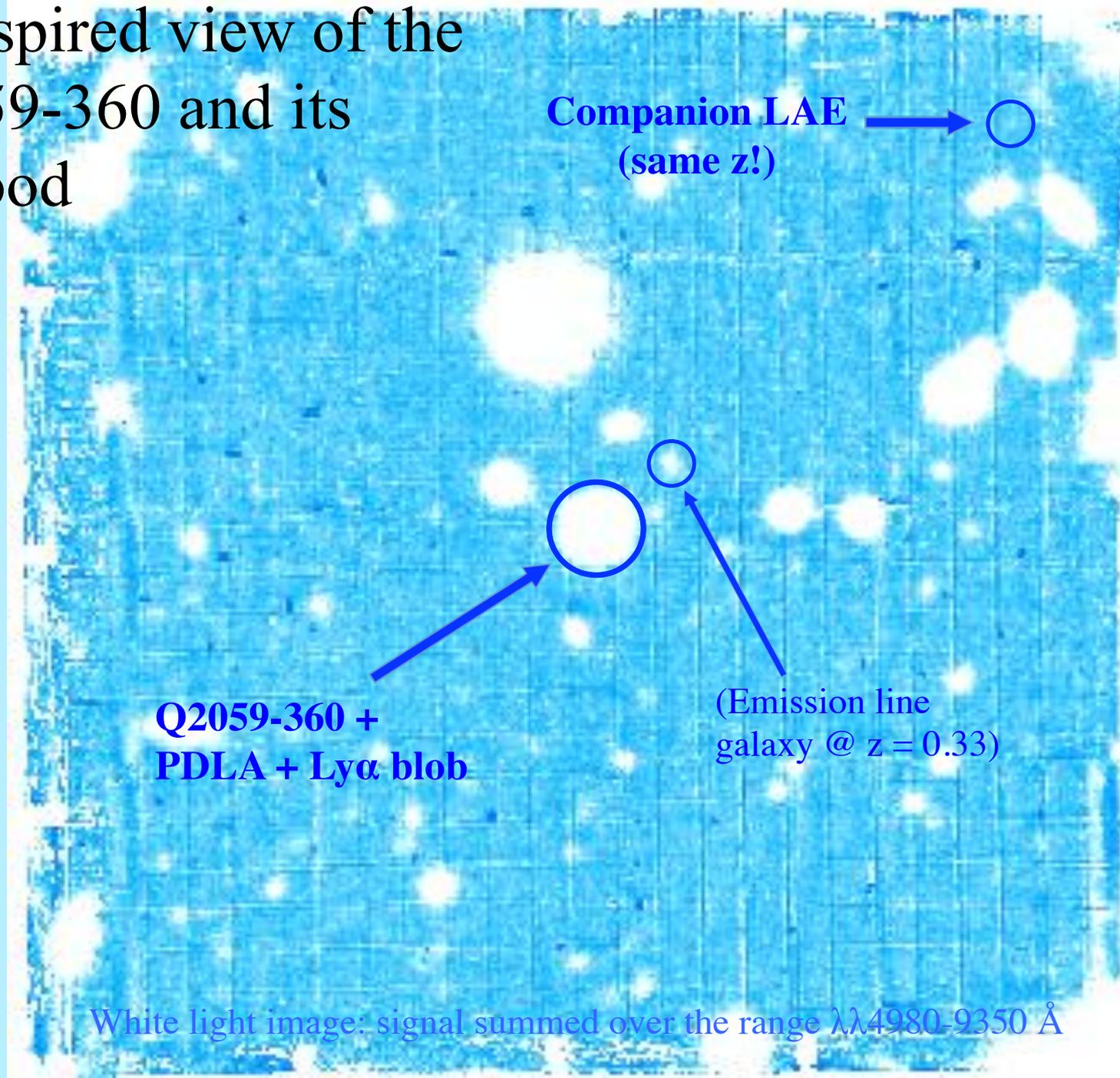


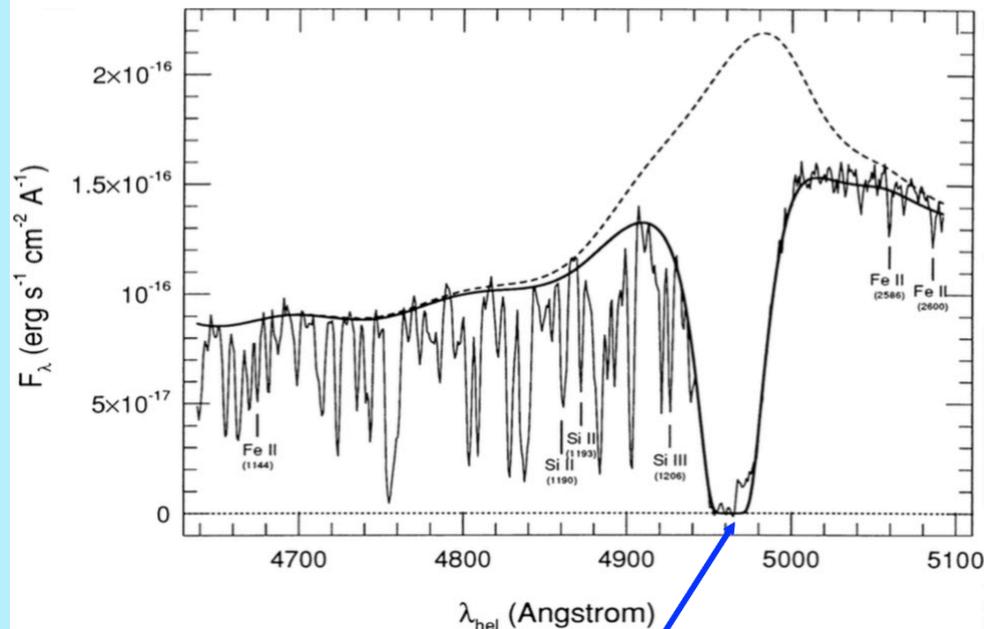
# A MUSE-inspired view of the quasar Q2059-360 and its neighbourhood

Pierre North  
Raffaella A. Marino  
Claudio Gorgoni  
Matthew Hayes  
Anne Verhamme  
Doron Chelouche  
Frédéric Courbin  
Dominique Sluse  
Sebastiano Cantalupo



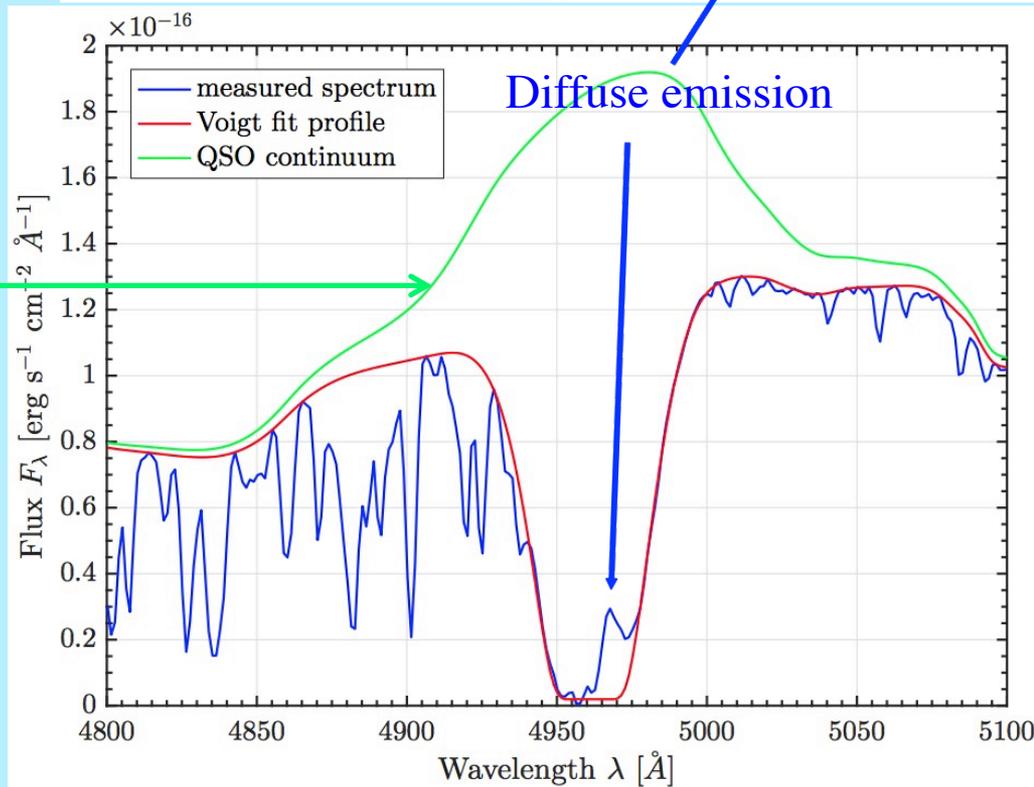
White light image: signal summed over the range  $\lambda\lambda 4980-9350 \text{ \AA}$

Leibundgut & Robertson 1999 (LR99):  
EMMI@NTT,  $t_{\text{exp}}=5$  hours, slit width 1.2''



MUSE@VLT,  $t_{\text{exp}}=3\text{h}20\text{mn}$ ,  
 $2.3 \times 2.8''$  area

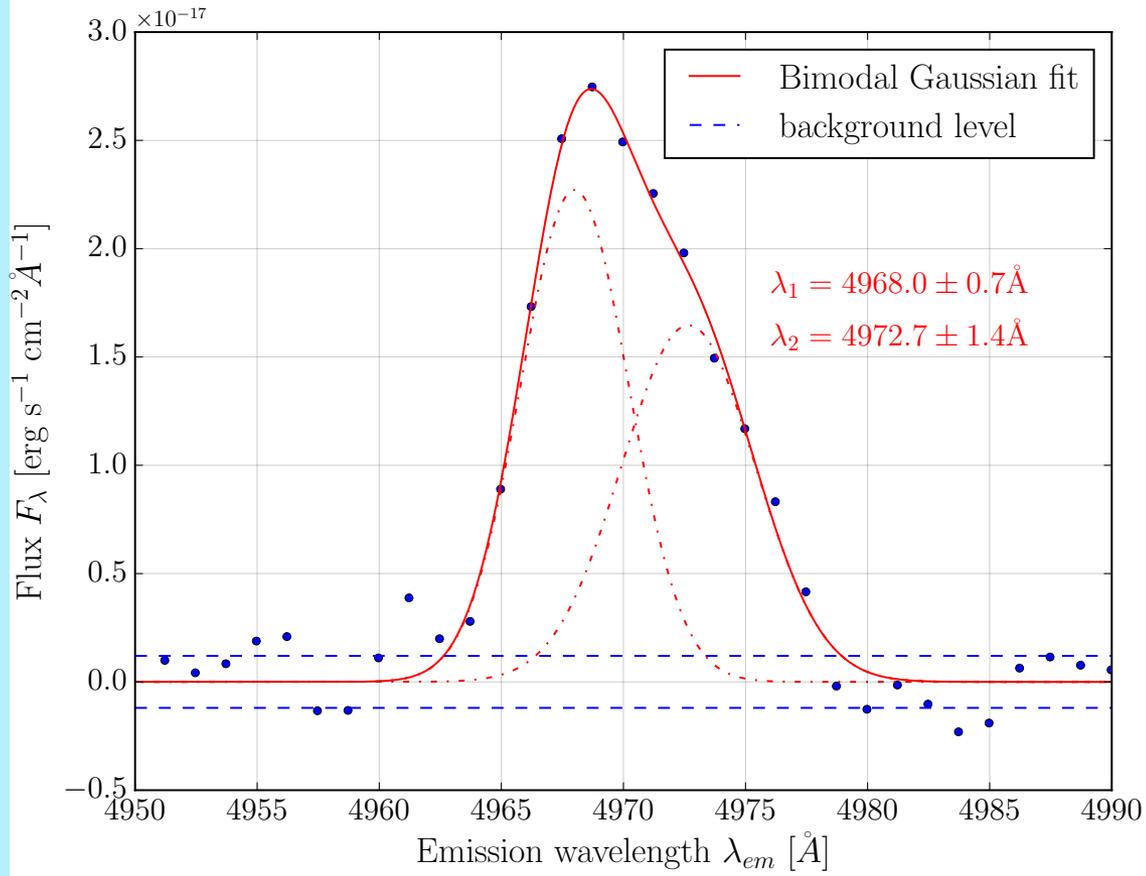
QSO « continuum » (actually broad Ly $\alpha$   
line) reconstructed using the wings of the  
DLA line (method of LR99)



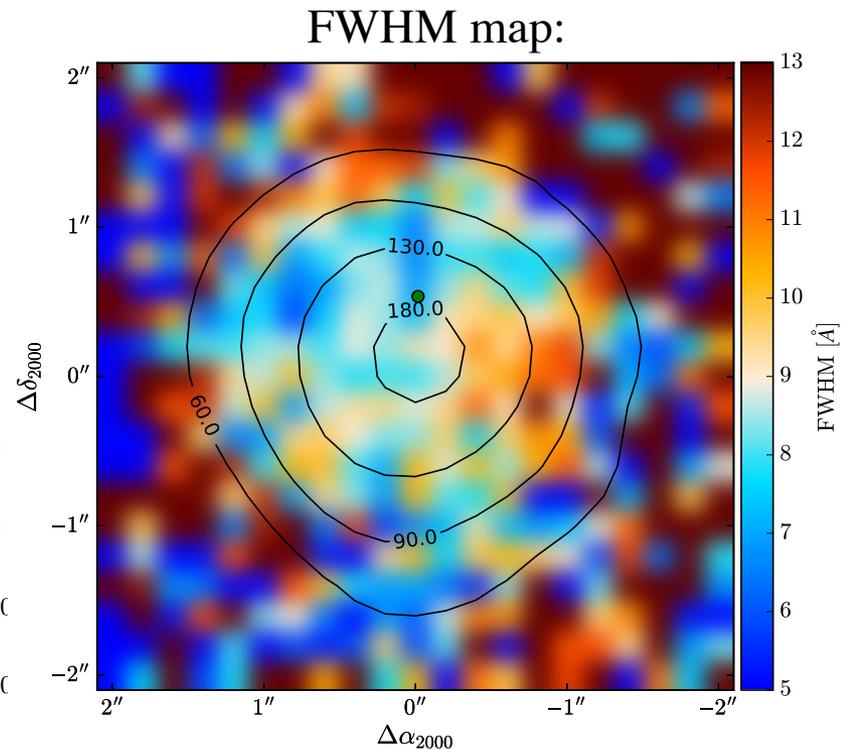
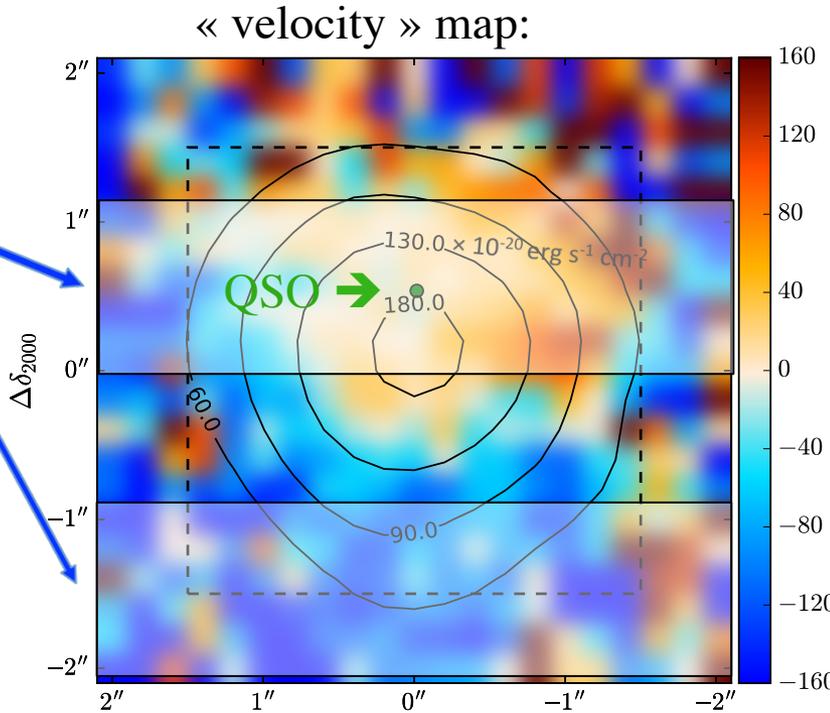
# Some properties of Q2059-360, its LAB and the PDLA:

Object	z	logN(HI)	F(Ly $\alpha$ )[erg/s/cm $^2$ ]	L(Ly $\alpha$ ) [erg/s]	$\Delta V$ [km/s]/ $\Delta\delta$ [ $''$ ]
QSO	3.09	-	-	-	-
PDLA	3.0818(1)	20.863(5)	-	-	-
LAB(core)	3.0885(2)	-	$(2.4 \pm 0.4) \times 10^{-16}$	$(2.1 \pm 0.4) \times 10^{43}$	$-60 \pm 10$ ( $\sim$ LR99)
Comp. LAE	3.0884(4) (main comp.)	-	$(8.9 \pm 1.7) \times 10^{-17}$	$(7.7 \pm 1.5) \times 10^{43}$	-

Integrated  
over a  
whole  
2.8'' $\times$  2.3''  
elliptical  
area

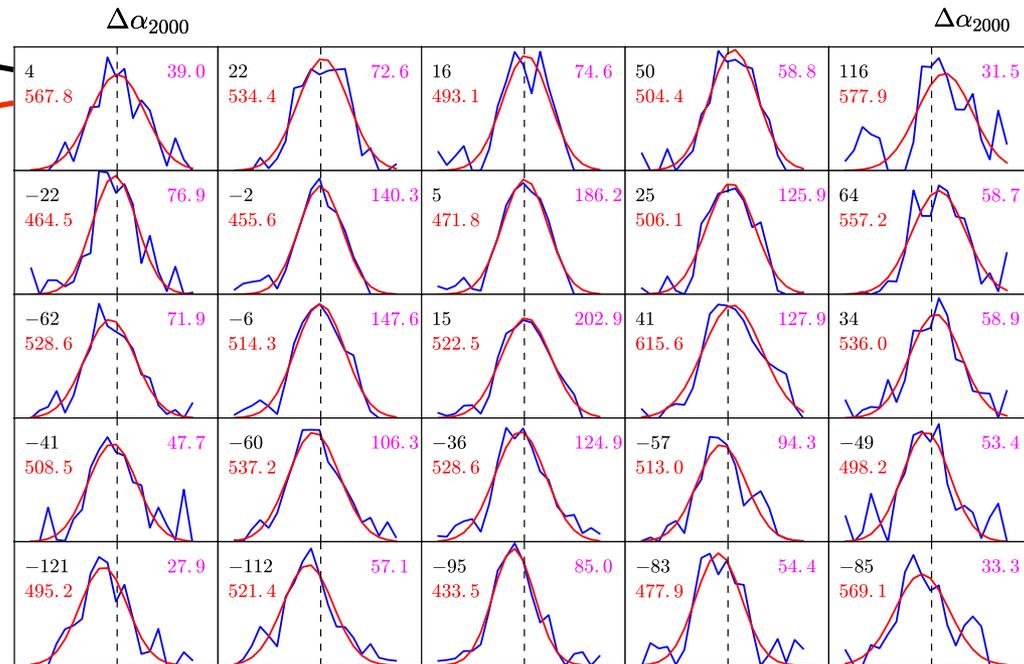


2 LR99  
slit  
positions



« velocity » [km/s]

FWHM [km/s]

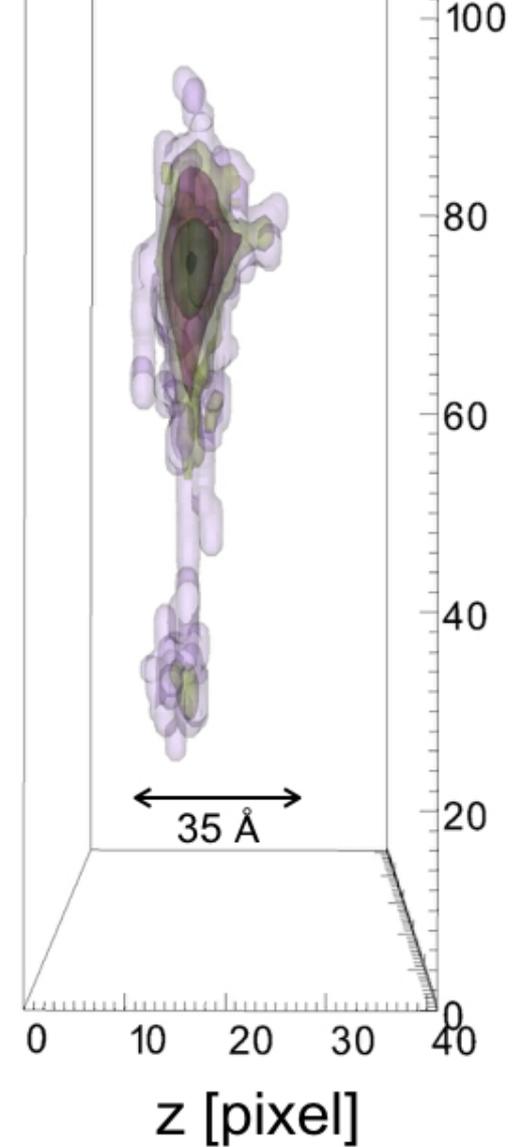
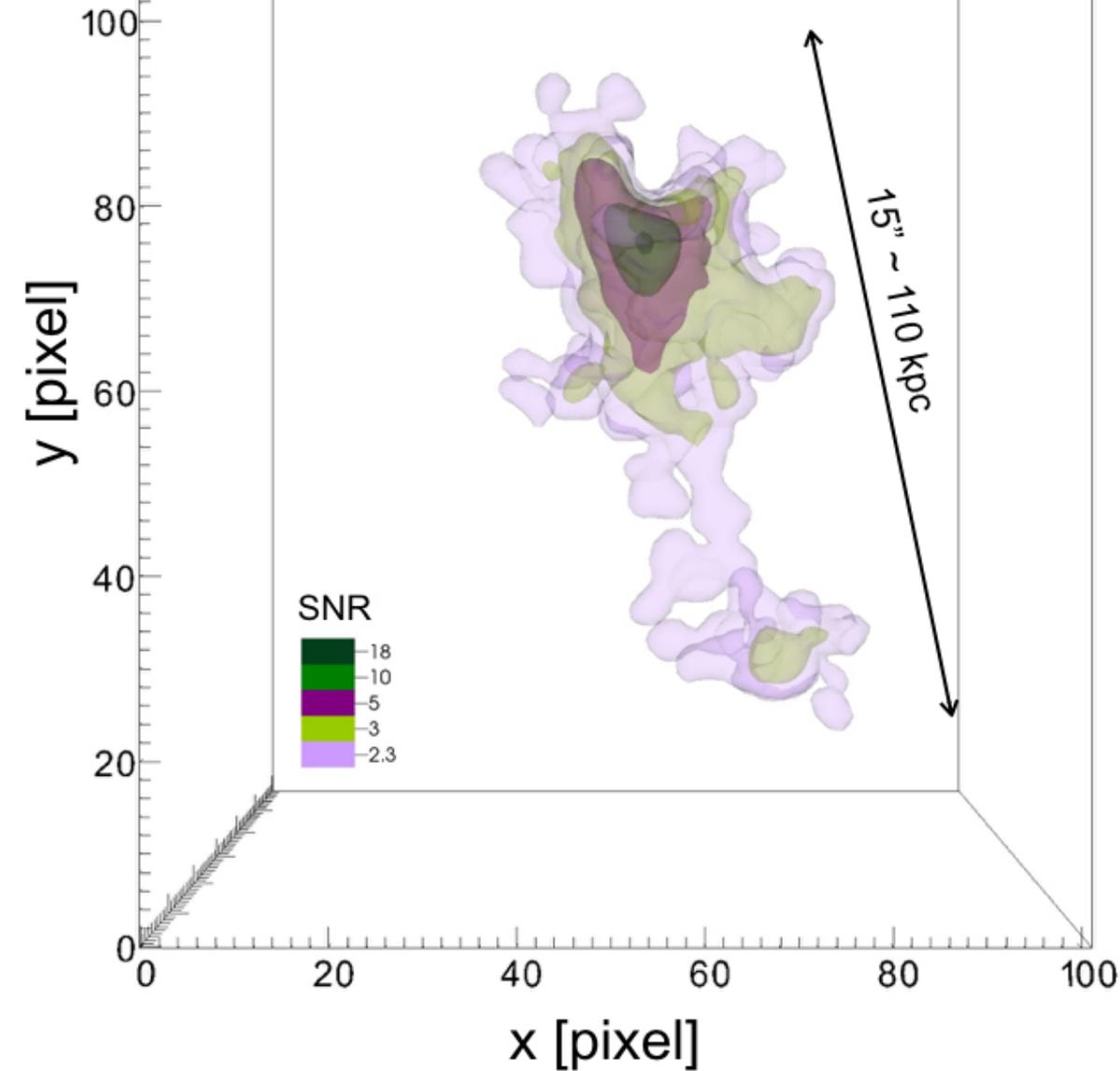


Each panel refers to  
a  $3 \times 3$  spaxels area  
( $0.6'' \times 0.6''$ )

**FWHM  $\leq \sim 600$  km/s  
 $\Rightarrow$  no violent outflow  
or turbulence**

The filament protruding from the main LAB:  $\sim 20$  kpc wide, and dynamically quiet (hardly resolved spectrally)

3D extraction (CubeExtractor, Cantalupo et al. in prep.)

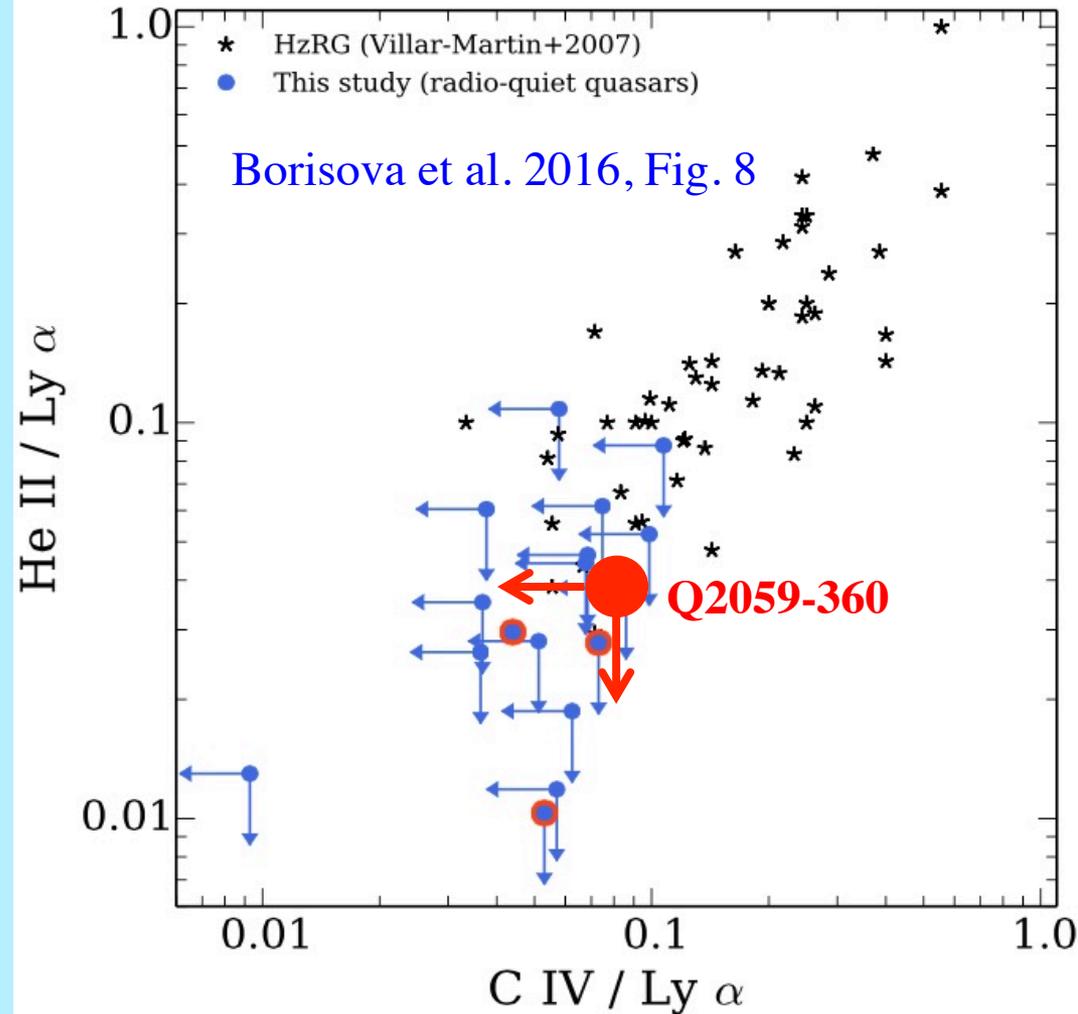


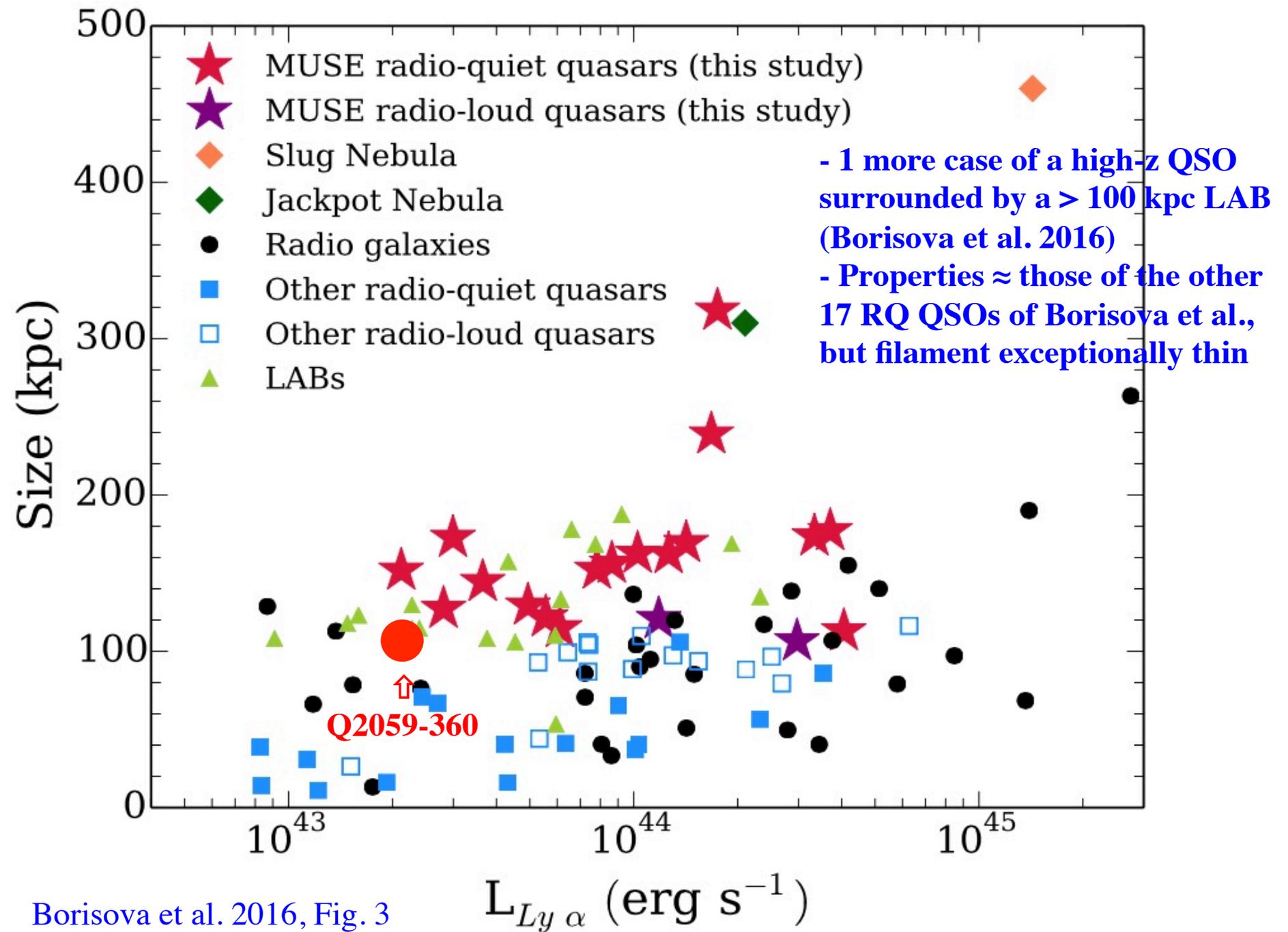
Upper limits on the N V, C IV, He II and C III] lines:

Line, wavelength (vac.)	Flux [ $\text{erg s}^{-1}\text{cm}^{-2}\text{arcsec}^{-2}$ ]	Flux(Line)/Flux(Ly $\alpha$ )
N V 1238.8-1242.8	$< 2.3 \times 10^{-18}$	$< 0.08$
C IV 1548.2-1550.8	$< 2.3 \times 10^{-18}$	$< 0.08$
He II 1640.4	$< 1.2 \times 10^{-18}$	$< 0.04$
C III] 1908.7	$< 3.5 \times 10^{-18}$	$< 0.12$

Limits on the average surface brightness measured in a disk with  $r = 4$  pixels, i.e. on a surface of  $\sim 2 \text{ arcsec}^2$ , with the disk centre  $1''$  south of the QSO, and assuming a 500 km/s FWHM (visual estimate).

The LAB surrounding Q2050-360 clearly differs from most HzRGs, while it appears similar to the 17 other RQ quasars examined by Borisova et al. 2016.

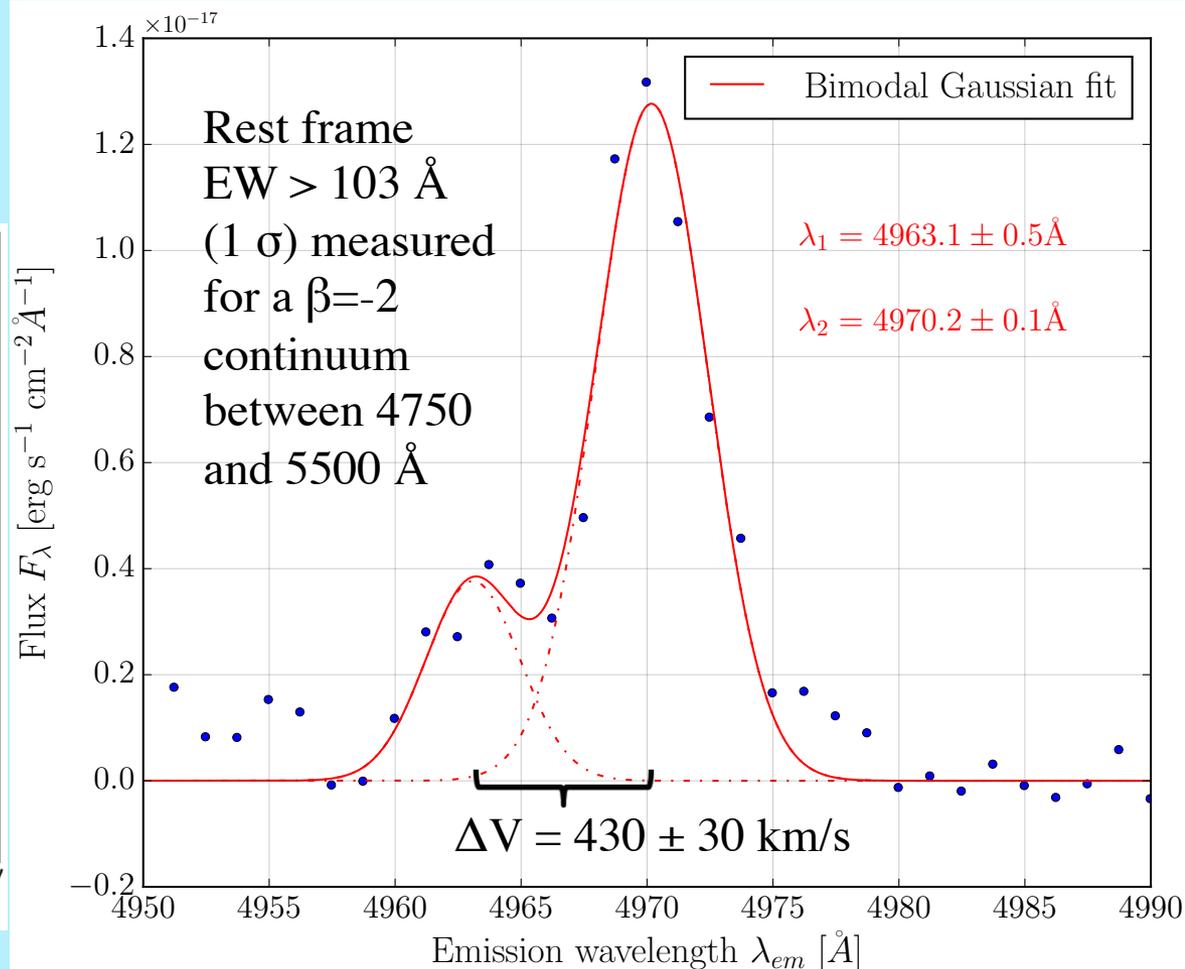
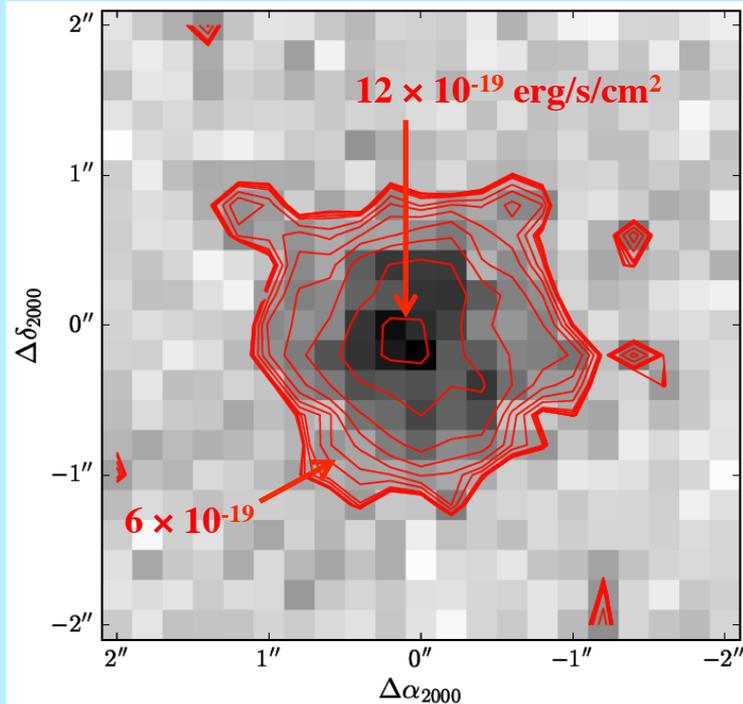




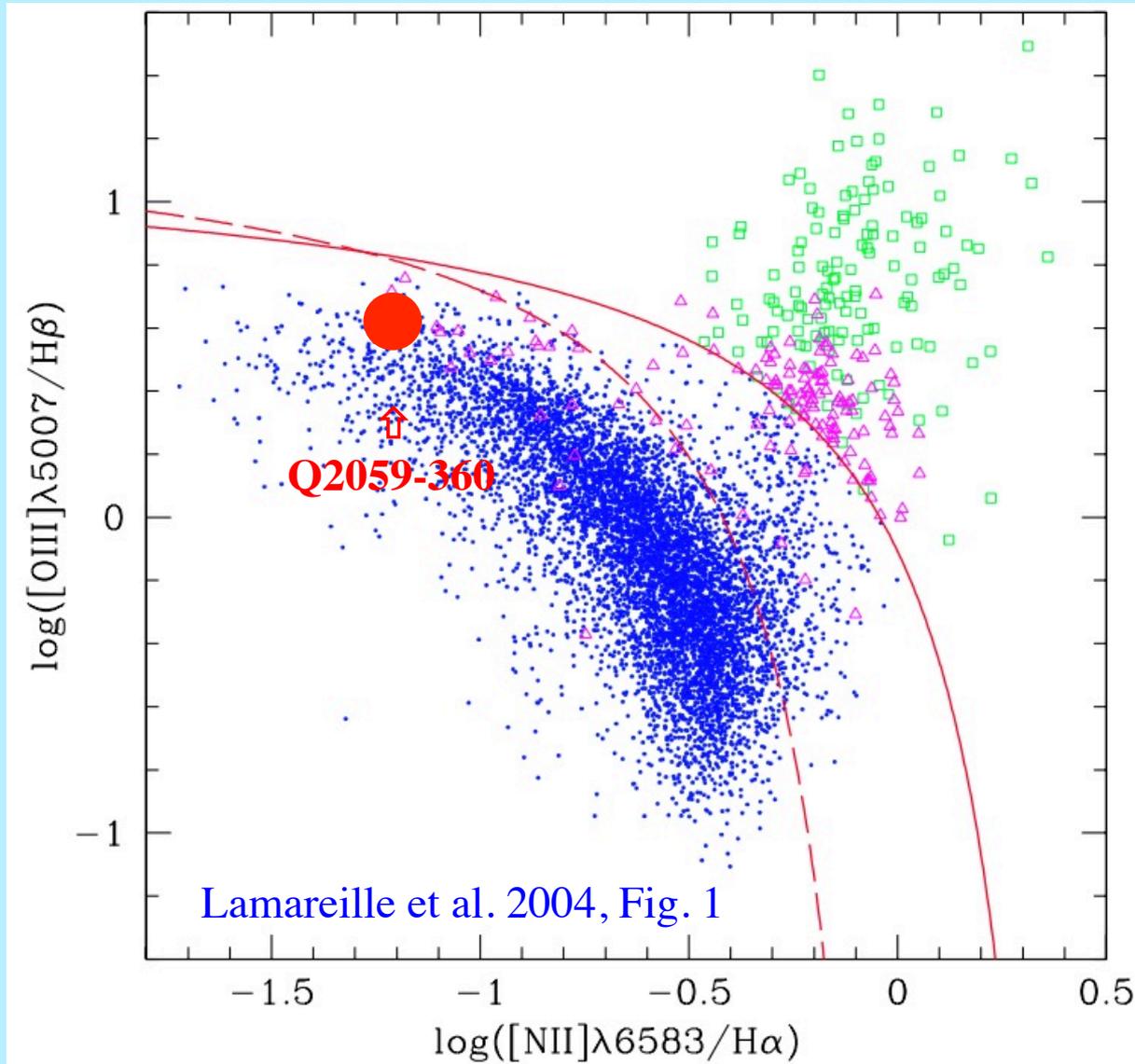
# The companion LAE at 34'' (265 kpc) to the NW:

$z$	$F(\text{Ly}\alpha)$ [erg/s/cm <sup>2</sup> ]	$L(\text{Ly}\alpha)$ [erg/s]	FWHM[km/s]
$3.0826 \pm 0.0004$ (blue comp.)	$(8.9 \pm 1.7) \times 10^{-17}$	$(7.7 \pm 1.5) \times 10^{43}$	$267 \pm 68$
$3.0884 \pm 0.0001$ (red comp.)			$317 \pm 23$

This object presumably shines in Ly $\alpha$ ; it is hardly resolved (seeing  $\sim 1.1''$ ):



The emission line galaxy at  $\alpha_{2000}=21:02:44.316$ ,  $\delta_{2000}=-35:53:02.55$ :  
a starburst galaxy



- Blue dots: star forming galaxies
- green squares: AGNs
- Magenta triangles: galaxies with discrepant classification ( $\text{OIII}5007/\text{H}\beta$  vs  $\text{SII}6717+6731/\text{H}\alpha$  diag.)